

# Relativistic fluctuation relation

Lecture 3 – The mathematical basis of relativity

**Lucas Chibebe Céleri**

Institute of Physics  
Federal University of Goiás

2025 – International Institute of Physics – Natal





# Differential Geometry



# The postulates of GR

General relativity is based on two postulates:

- Spacetime is a four-dimensional differentiable manifold with a Lorentzian metric defined on it.
- Einstein's field equations hold

$$R_{ab} - \frac{1}{2}Rg_{ab} + \lambda g_{ab} = 8\pi GT_{ab}$$

We need differential geometry in order to properly understand these postulates.

# Topological spaces



Let  $\mathbb{X}$  be a set and  $\mathbb{T}$  a collection of subsets of  $\mathbb{X}$  satisfying the following properties.

1. The union of an arbitrary collection of subsets, each of which is in  $\mathbb{T}$ , is in  $\mathbb{T}$ . If  $O_\alpha \in \mathbb{T}$  for all  $\alpha$ , then  $\cup_\alpha O_\alpha \in \mathbb{T}$ .
2. The intersection of a finite number of subsets of  $\mathbb{T}$  is in  $\mathbb{T}$ . If  $\{O_\alpha\}_{\alpha=1}^n \in \mathbb{T}$ , then  $\cap_{\alpha=1}^n O_\alpha \in \mathbb{T}$ .
3. The entire set  $\mathbb{X}$  and the empty set  $\emptyset$  are in  $\mathbb{T}$ .

$\mathbb{X}$  is said to be a topological space, and  $\mathbb{T}$  provides a topology to  $\mathbb{X}$ .

# Examples



- The real line  $\mathbb{R}$  where the open sets are unions of open intervals  $(a, b) \subset \mathbb{R}$ .
- $\mathbb{X} = \{a, b, c\}$  and  $\mathbb{T} = \{\mathbb{X}, \emptyset, \{a, b\}, \{b\}, \{b, c\}\}$ . The pair  $(\mathbb{X}, \mathbb{T})$  is a topological space. The collection of all subsets of  $\mathbb{X}$  is called the discrete topology, while the set  $\{\mathbb{X}, \emptyset\}$  is called the trivial topology.
- If a metric<sup>1</sup>  $d(x, y)$ , with  $x, y \in \mathbb{X}$ , is defined on  $\mathbb{X}$ , its open sets are given by the open discs  $\mathbb{O}_\epsilon = \{y \in \mathbb{X} \mid d(x, y) < \epsilon\}$  and all their possible unions. Such topology is called a metric topology determined by  $d$  and the topological space  $(\mathbb{X}, \mathbb{T})$  is called a **metric space**.

---

<sup>1</sup> $d : \mathbb{X} \times \mathbb{X} \mapsto \mathbb{R}$  under the conditions *i*)  $d(x, y) = d(y, x)$ , *ii*)  $d(x, y) \geq 0$ , with the equality holding only for  $x = y$  and *iii*)  $d(x, y) + d(y, z) \geq d(x, z)$  for all  $x, y, z \in \mathbb{X}$ .

# What is a manifold?



## Manifold

is just a continuous set of points that can exhibit complex global properties like curvature or torsion, but that *locally* looks like the Euclidean space. In a small enough neighborhood of any point on the manifold, Euclidean geometry applies. The surface of a sphere is a manifold!.

## The meaning of locally flat space

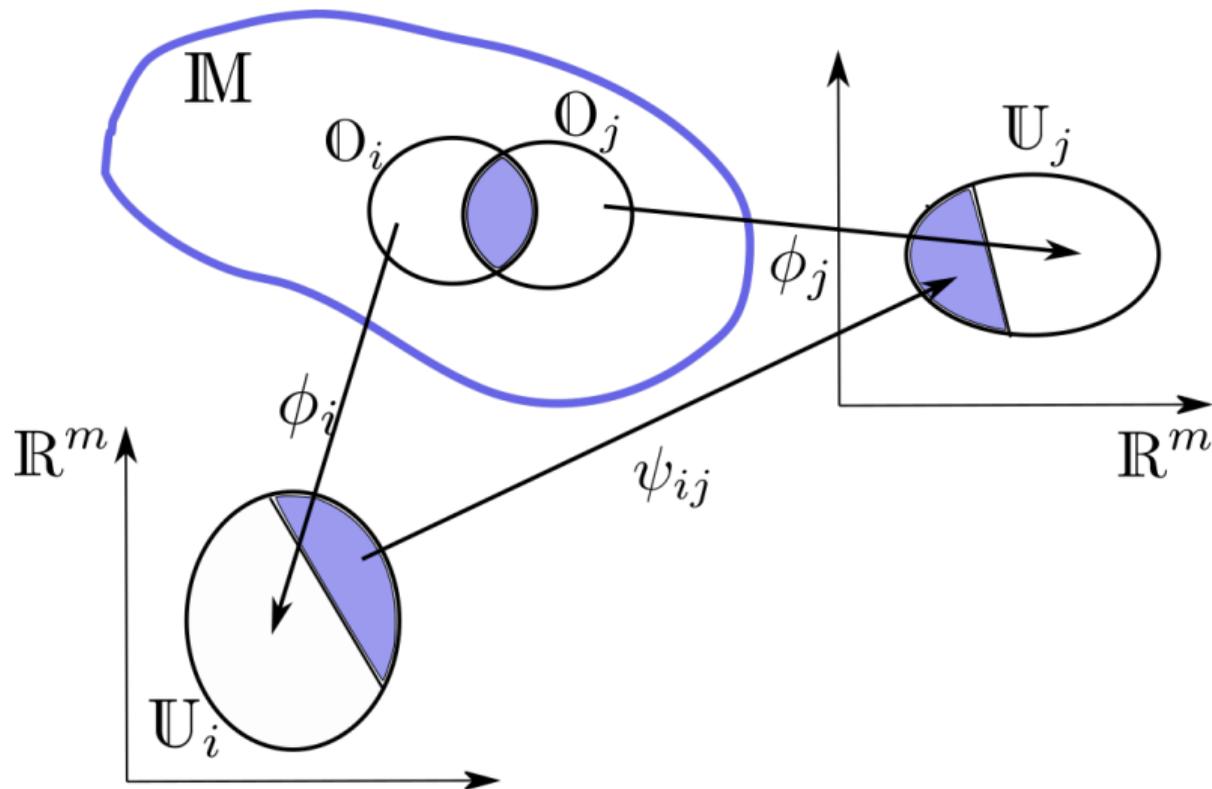
A manifold is a generalization of the concepts of curves and surfaces to objects with arbitrary dimensions. In the same way a curve is locally homeomorphic to  $\mathbb{R}$  a manifold is a topological space which is locally homeomorphic to  $\mathbb{R}^m$ . This is what enables us to define local coordinate systems!

## Formal definition

$\mathbb{M}$  is an  $m$ -dimensional differentiable manifold if we have a set of pairs  $\{(\mathcal{O}_i, \phi_i)\}_i$  of open subsets (local neighborhoods)  $\mathcal{O}_i \subset \mathbb{M}$  and local homeomorphisms  $\phi_i$  from  $\mathcal{O}_i$  to open subsets  $\mathcal{U}_i \subset \mathbb{R}^m$ , such that the following conditions hold:

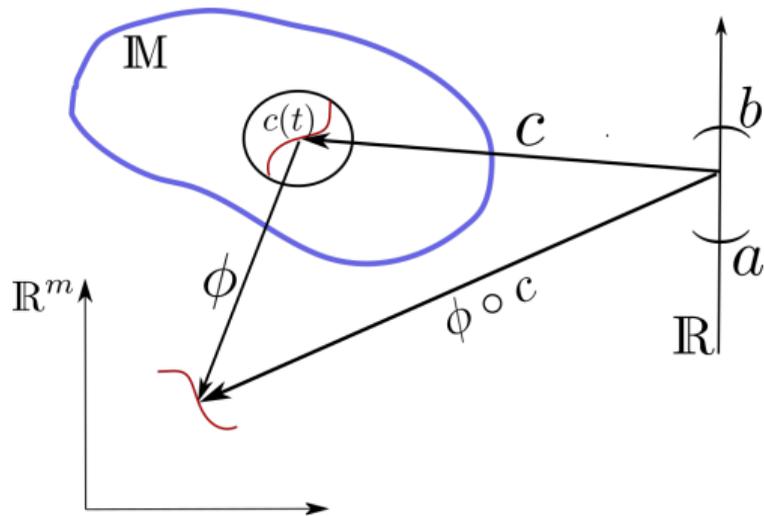
- $\mathbb{M}$  is a topological space.
- The set of  $\mathcal{O}_i$  covers  $\mathbb{M}$ :  $\cup_i \mathcal{O}_i = \mathbb{M}$ .
- Given two neighbourhoods  $\mathcal{O}_i$  and  $\mathcal{O}_j$  such that  $\mathcal{O}_i \cap \mathcal{O}_j \neq \emptyset$  for  $i \neq j$ , the map  $\psi_{ij} = \phi_i \circ \phi_j^{-1} : \phi_j(\mathcal{O}_i \cap \mathcal{O}_j) \mapsto \phi_i(\mathcal{O}_i \cap \mathcal{O}_j)$  is infinitely differentiable.

# Differentiable structure



# Curves

An open curve on  $M$  is the map  $c : (a, b) \mapsto M$ , where  $(a, b) \subset \mathbb{R}$ .  $c(t)$  has the coordinate representation  $x = \phi \circ c : \mathbb{R} \mapsto \mathbb{R}^m$ .



# Vectors



## Tangent space

On a manifold, a vector is defined in terms of a **tangent vector** to a curve, which is the generalization of the tangent line to a curve in the usual two-dimensional plane. However, in an  $n$ -dimensional manifold, there are many curves that are indeed tangent to a given vector and, therefore, a tangent vector is an equivalence class of curves.

# Vectors



Let us consider a function  $f : M \mapsto \mathbb{R}$ . In a local coordinate system

$$\left. \frac{df(c(t))}{dt} \right|_{t=0} = \left. \frac{\partial f}{\partial x^\mu} \frac{dx^\mu(c(t))}{dt} \right|_{t=0} \quad \frac{\partial f}{\partial x^\mu} \equiv \frac{\partial}{\partial x^\mu} [f \circ \phi^{-1}(x)]$$

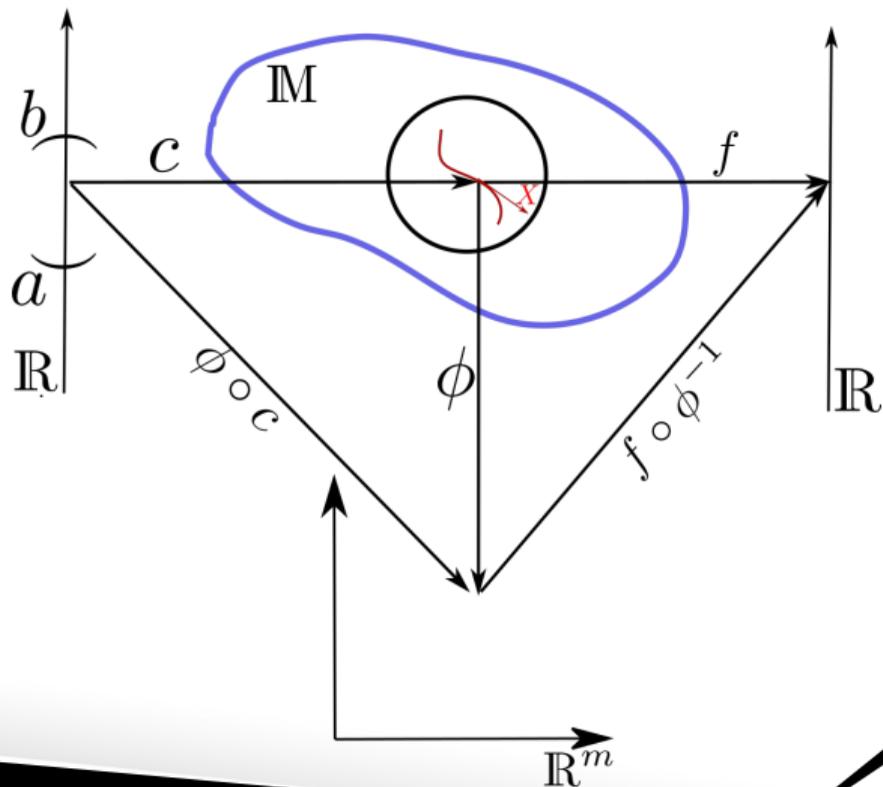
The rate  $(df/dt)|_{t=0}$  is obtained by the application of the differential operator

$$X = X^\mu \left( \frac{\partial}{\partial x^\mu} \right) \quad \text{with} \quad X^\mu = \left. \frac{dx^\mu(c(t))}{dt} \right|_{t=0}$$

to the map  $f$ , denoted as  $X[f]$ .

# Vectors

We define  $X$  as the tangent vector to  $\mathbb{M}$  at the point  $p = c(0)$  along the direction given by the curve  $c(t)$ .



# The tangent space

Given two curves  $c_1$  and  $c_2$  such that  $c_1(0) = c_2(0) = p$  and

$$\left. \frac{dx^\mu(c_1(t))}{dt} \right|_{t=0} = \left. \frac{dx^\mu(c_2(t))}{dt} \right|_{t=0}$$

then both curves define the same vector  $X$  at  $p$  ( $c_1 \sim c_2$ ).  $X$  is the equivalence class of curves

$$[c(t)] = \left\{ \tilde{c}(t) \mid \tilde{c}(0) = c(0) \quad \text{and} \quad \left. \frac{dx^\mu(c_1(t))}{dt} \right|_{t=0} = \left. \frac{dx^\mu(c_2(t))}{dt} \right|_{t=0} \right\},$$

All the equivalence classes of curves at  $p \in \mathbb{M}$  (all the tangent vectors at  $p$ ) form a vector space called the **tangent space** of  $\mathbb{M}$  at  $p$ , denoted by  $T_p\mathbb{M}$ .



# The coordinate basis

We have

$$\frac{\partial x^\mu}{\partial x^\nu} = \delta_\nu^\mu \Rightarrow X[x^\mu] = \left. \frac{dx^\mu(t)}{dt} \right|_{t=0}$$

The coordinate basis is then defined by the vectors

$$e_\mu \equiv \frac{\partial}{\partial x^\mu}$$

We have a tangent space associated with each point on the manifold. The collection of all these tangent spaces for all points on the manifold is called the **tangent bundle**, which is a manifold in its own right.



## Transformation laws

The vector is coordinate independent, but its components are not. Let  $p \in \mathbb{O}_i \cap \mathbb{O}_j$  and  $\{x^\mu\} = \phi_i(p)$  and  $\{y^\nu\} = \phi_j(p)$  be the two coordinate functions. We then have two expressions for the vector  $X \in T_p\mathbb{M}$

$$X = X^\mu \frac{\partial}{\partial x^\mu} = \tilde{X}^\mu \frac{\partial}{\partial y^\mu}.$$

This implies that the components of the vector in the two coordinate bases must be related by

$$\tilde{X}^\mu = X^\nu \frac{\partial y^\mu}{\partial x^\nu}.$$

It is important to observe here that this transformation law is such that the vector itself is left invariant.

# The dual space



We now associate the **dual vector space** (cotangent space)  $T_p^*\mathbb{M}$  to every  $T_p\mathbb{M}$ .

$$\omega \in T_p^*\mathbb{M} : T_p\mathbb{M} \mapsto \mathbb{R}$$

$\omega$  is called the cotangent vector, the dual vector, or the one-form, with the simplest example being the differential  $df$  of a function  $f \in \mathbb{F}(\mathbb{M})$ . The action of  $df \in T_p^*\mathbb{M}$  on  $V \in T_p\mathbb{M}$  is defined as

$$\langle df, V \rangle = V[f] = V^\mu \frac{\partial f}{\partial x^\mu} \in \mathbb{R}$$

Since  $df = (\partial f / \partial x^\mu) dx^\mu$ , we take  $dx^\mu$  as the elements of the basis in  $T_p^*\mathbb{M}$ , which is the dual basis since

$$\left\langle dx^\nu, \frac{\partial}{\partial x^\mu} \right\rangle = \frac{\partial x^\nu}{\partial x^\mu} = \delta_\mu^\nu \quad \Rightarrow \quad \omega = \omega_\mu dx^\mu$$

# The inner product

The inner product  $\langle \cdot, \cdot \rangle : T_p^* \mathcal{M} \times T_p \mathcal{M} \mapsto \mathbb{R}$  is defined as

$$\langle \omega, V \rangle = \omega_\nu V^\mu \left\langle dx^\nu, \frac{\partial}{\partial x^\mu} \right\rangle = \omega(V) = \omega_\mu V^\mu.$$

Let us consider two coordinate systems  $\{x^\mu\} = \phi_i(p)$  and  $\{y^\nu\} = \phi_j(p)$ . We thus have

$$\omega = \omega_\mu dx^\mu = \tilde{\omega}_\nu dy^\nu$$

From which follows the transformation law for the components of the one-form

$$\tilde{\omega}_\nu = \omega_\mu \frac{\partial x^\mu}{\partial y^\nu}$$

# Tensors



A tensor of type  $(q, r)$  is a multilinear map

$$\mathbb{T} : [\times^q \mathbb{T}_p^* \mathbb{M}] [\times^r \mathbb{T}_p \mathbb{M}] \mapsto \mathbb{R}$$

In terms of the basis vectors defined before, we can write the tensor  $\mathbb{T} \in \mathbb{T}_{r,p}^q(\mathbb{M})$  as

$$\mathbb{T} = T^{\mu_1 \dots \mu_q}{}_{\nu_1 \dots \nu_r} \frac{\partial}{\partial x^{\mu_1}} \dots \frac{\partial}{\partial x^{\mu_q}} dx^{\nu_1} \dots dx^{\nu_r}$$

The action of a tensor on  $\omega_i$  ( $1 \leq i \leq r$ ) and  $V_j$  ( $1 \leq j \leq q$ ) results in the number

$$\mathbb{T}(\omega_1, \dots, \omega_r; V_1, \dots, V_q) = T^{\mu_1 \dots \mu_q}{}_{\nu_1 \dots \nu_r} \omega_{1, \mu_1} \dots \omega_{q, \mu_q} V_1^{\nu_1} \dots V_r^{\nu_r}$$

# Tensors



Given two coordinate systems  $x$  and  $x'$ , the components of the tensor  $T$  change as

$$T^{\mu'_1 \dots \mu'_q}_{\nu'_1 \dots \nu'_r} = T^{\mu_1 \dots \mu_q}_{\nu_1 \dots \nu_r} \frac{dx^{\mu'_1}}{dx^{\mu_1}} \cdots \frac{dx^{\mu'_q}}{dx^{\mu_q}},$$

which is the general transformation law for tensors.

We can let the tensor not act on all of its slots, resulting in different tensors. For instance, a map between two vector spaces can be seen as a tensor

$$T^a_b : V^b \mapsto T^a_b V^b$$

# The metric tensor

We impose that the determinant of the metric does not vanish,  
 $\text{Det}[g_{ab}] \equiv g \neq 0$ . Then

$$g^{\mu\nu} g_{\alpha\nu} = \delta_{\alpha}^{\mu}.$$

One of the many applications of the metric tensor is to provide the notion of the length of a path in spacetime, the line element

$$ds^2 = g_{\mu\nu} dx^{\mu} \otimes dx^{\nu} = g_{\mu\nu} dx^{\mu} dx^{\nu}.$$

The metric is employed to raise and lower indices

$$x_a = g_{ab} x^b \qquad x^a = g^{ab} x_b$$



# The covariant derivative

Vectors and tensors are coordinate-independent objects. The laws of physics must be written in terms of them!

But we need derivatives. The partial derivative is not a tensor since it does not transform like a tensor does. If we want the laws of physics to be coordinate-independent, we need to define the notion of the covariant derivative on manifolds.

From this, the important concept of a **connection** emerges. The curvature of a manifold depends on this guy!

# The covariant derivative

In flat space in inertial coordinates, the partial derivative operator is a map between tensor fields

$$\partial_\mu : (k, l) \rightarrow (k, l + 1)$$

that obeys linearity and the Leibniz rule.

We would like to define an operator  $\nabla$  that plays the role of the partial derivative but that is coordinate independent! We require that  $\nabla : (k, l) \rightarrow (k, l + 1)$  that obeys:

- **Linearity:**  $\nabla(T + S) = \nabla T + \nabla S$
- **Leibniz:**  $\nabla(T \otimes S) = \nabla(T) \otimes S + T \otimes \nabla(S)$

## The connection

From Leibniz's rule,  $\nabla$  is written as a partial derivative and a linear correction to make it covariant. There must be a set of  $n$  matrices  $(\Gamma_\mu)^\alpha_\beta$  for each direction  $\mu$  such that

$$\nabla_\mu V^\nu = \partial_\mu V^\nu + \Gamma_{\mu\alpha}^\nu V^\alpha$$

$\Gamma_{\mu\alpha}^\nu$  are the **connection coefficients**. By demanding that it commutes with contraction and that it reduces to the partial derivative in flat spacetimes, we obtain

$$\nabla_\mu \omega_\nu = \partial_\mu \omega_\nu - \Gamma_{\mu\nu}^\alpha \omega_\alpha$$

From this we can construct the derivative of higher tensors.

# The compatible connection

There are many distinct connections! Which one should we pick for GR?

Although  $\Gamma_{\mu\nu}^{\alpha}$  is not a tensor, the difference between two of them is the torsion tensor

$$S^{\alpha}_{\beta\gamma} = \Gamma_{\beta\gamma}^{\alpha} - \Gamma_{\gamma\beta}^{\alpha}$$

A unique connection arises in GR by imposing:

- Torsion free:  $S^{\alpha}_{\beta\gamma} = 0$
- Metric compatibility:  $\nabla_{\mu}g_{\alpha\beta} = 0$

From which the Christoffel connection follows

$$\Gamma_{\beta\gamma}^{\alpha} = \frac{1}{2}g^{\alpha\sigma} (\partial_{\beta}g_{\gamma\sigma} + \partial_{\gamma}g_{\beta\sigma} - \partial_{\sigma}g_{\beta\gamma})$$



## What do we have done so far?

- We started with the basic notion of a set.
- We introduced the concept of open subsets. The set became a topological space.
- By demanding that each open set look like a region of  $\mathbb{R}^n$  and that the coordinate charts be smoothly sewn together, the topological space became a manifold.
- A manifold comes equipped naturally with a tangent bundle, tensor bundles, the ability to take exterior derivatives, and so forth.
- We then proceeded to put a metric on the manifold.
- Independently of the metric we could introduce a connection, allowing us to take covariant derivatives.

# How to compare vectors from different spaces?



We need to understand how to move vectors along paths while *keeping them constant*. This is the job of **parallel transport**, which is defined by the connection!

## Path dependence

The key difference between transporting a vector in flat and curved spacetimes is that it is path-dependent in the curved case!

There is no way to define relative velocities for events happening at distinct points of spacetime!

# The geodesic equation

The **geodesic** is defined as the curve along which the tangent vector is parallel transported (in the same way the tangent vector is transported along a straight line!)

$$\frac{D}{d\lambda} \frac{dx^\mu(\lambda)}{d\lambda} = 0$$

This implies the **geodesic equation**

$$\frac{d^2 x^\alpha(\lambda)}{d\lambda^2} + \Gamma_{\mu\nu}^\alpha \frac{dx^\mu(\lambda)}{d\lambda} \frac{dx^\nu(\lambda)}{d\lambda} = 0$$

Extremals of the length functional are curves that parallel transport their tangent vector with respect to the Christoffel connection associated with that metric.



# The curvature

The curvature of a manifold is quantified by the Riemann curvature tensor, which is derived from the connection!

## Manifestations of flatness

- Parallel transport around a closed loop leaves a vector unchanged.
- Covariant derivatives of tensors commute.
- Initially parallel geodesics remain parallel.

The Riemann tensor arises when we study how any of these properties are altered in more general contexts!

## The curvature

Let us consider an infinitesimal closed loop specified by the vectors  $A^\mu$  and  $B^\nu$ .

A vector  $V^\mu$  is first parallel transported along  $A^\mu$ , then along  $B^\nu$ , then backward along  $A^\mu$  and  $B^\nu$ .

Parallel transport is independent of coordinates, so there must be a tensor telling us how the vector changes when it comes back to its starting point. It will be a linear transformation on a vector, and therefore involve one upper and one lower index. It also depends on the two vectors  $A^\mu$  and  $B^\nu$ , implying two lower indices. The change  $\delta V^\sigma$  must be written as

$$\delta V^\sigma = R^\sigma_{\alpha\beta\gamma} V^\alpha A^\beta B^\gamma$$

# The curvature



This is the definition of the Riemann tensor  $R^\sigma_{\alpha\beta\gamma} V^\alpha A^\beta B^\gamma$ . However, since we know how covariant derivatives work, we can compute the commutator

$$[\nabla_\mu, \nabla_\nu]V^\alpha = \nabla_\mu \nabla_\nu V^\alpha - \nabla_\nu \nabla_\mu V^\alpha$$

From which we obtain the formula for the curvature tensor in terms of the connection

$$R^\rho_{\sigma\mu\nu} = \partial_\mu \Gamma^\rho_{\nu\sigma} - \partial_\nu \Gamma^\rho_{\mu\sigma} + \Gamma^\rho_{\mu\lambda} \Gamma^\lambda_{\nu\sigma} - \Gamma^\rho_{\nu\lambda} \Gamma^\lambda_{\mu\sigma}$$

From which it follows three important quantities

$$R_{\mu\nu} = R^\rho_{\mu\rho\nu} \quad R = R^\mu_{\mu} = g^{\mu\nu} R_{\mu\nu} \quad G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu}$$

# That is all, Folks!

[lucas@qpequi.com](mailto:lucas@qpequi.com)

[www.qpequi.com](http://www.qpequi.com)

